Investigating Stars: Exploration Two: Activity 1

Stellar Classes and Wavelengths of Important Elements

Spectral Type	Temperature (K)	Spectral Lines
0	28 — 50,000	Helium, very few lines overall, no hydrogen lines
В	10-28,000	A little helium, some hydrogen
А	7,500 — 10,000	Very strong hydrogen, some sodium
F	6,000 — 7,500	Weaker hydrogen, stronger sodium, weak calcium lines, iron visible but weaker than $H_{_{gamma}}$
G	5 — 6,000	Weak hydrogen, strong iron, calcium lines very strong, iron strong
К	3.500 — 5,000	Many lines visible between $H_{_{gamma}}$ and $H_{_{\Delta'}}$ iron weak, hydrogen almost invisible, calcium weaker, iron weaker
М	2,500 — 3,500	Titanium oxide bands at 4750 and 4950 Å, many, many lines in spectrum

Wavelengths of Some Elements in Stellar Spectra

Spectral Lines	Wavelengths (Å) (rounded off)
Hydrogen $(H_{\sigma'} H_{\beta'} H_{gamma'} H_{\Delta})$	6600, 4800, 4350, 4100
Calcium	3930, 3960
Helium	4200, 4400
Sodium	5800
Iron	4300
Titanium Oxide bands	Lots of lines between 4900 and 5200, 5400-5700, 6200-6900